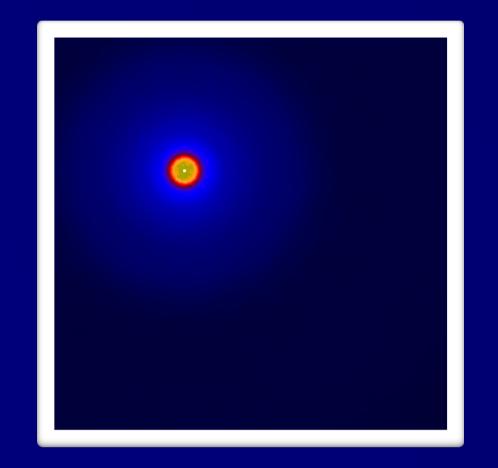
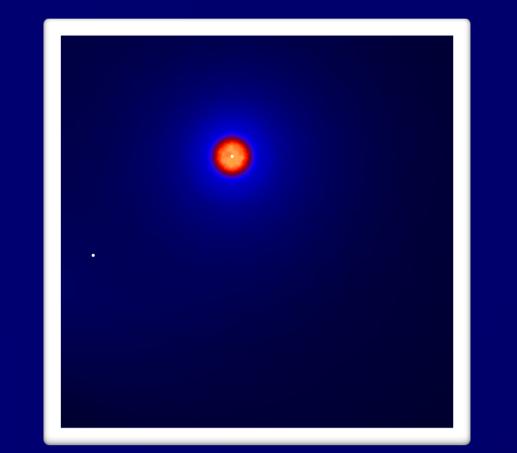
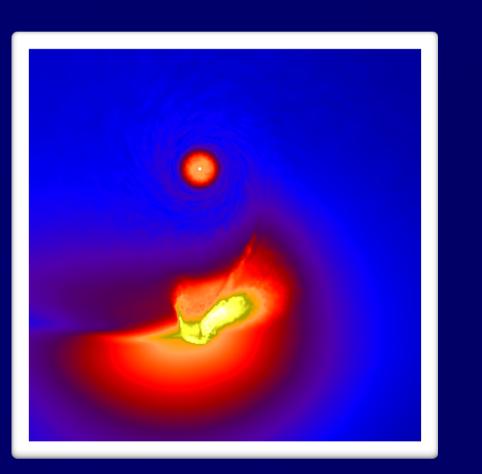
## **Disc-penetrating** stellar flyby

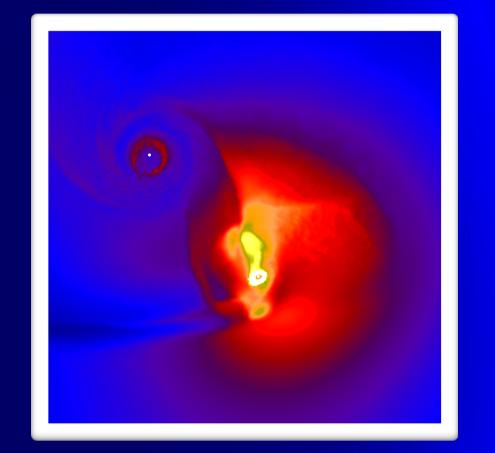
## \*triggers fast FU Orionis outburst

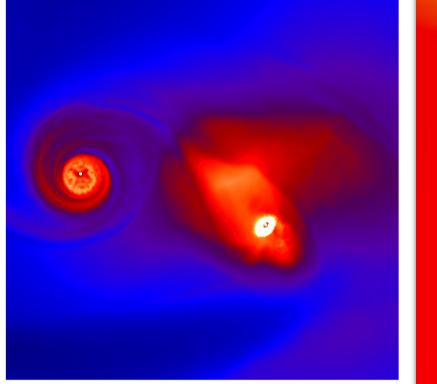
## on perturbing star

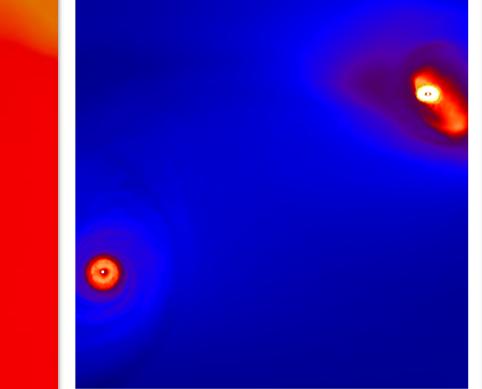


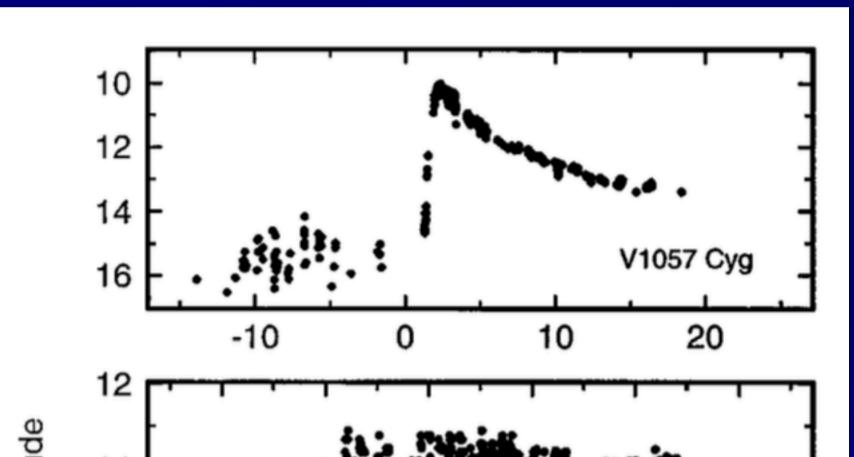




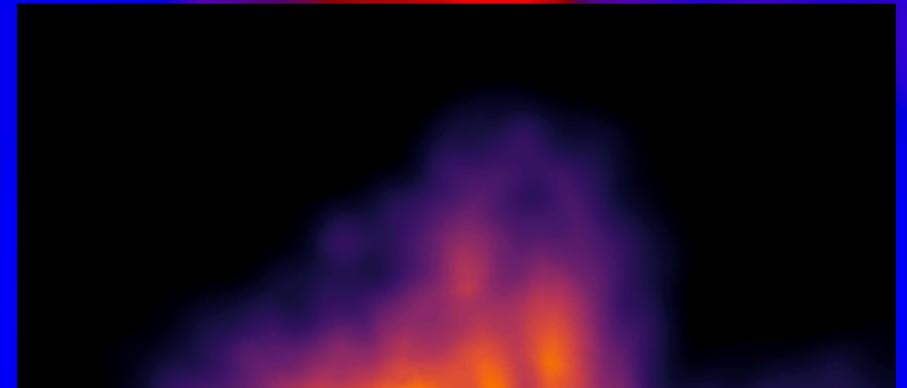


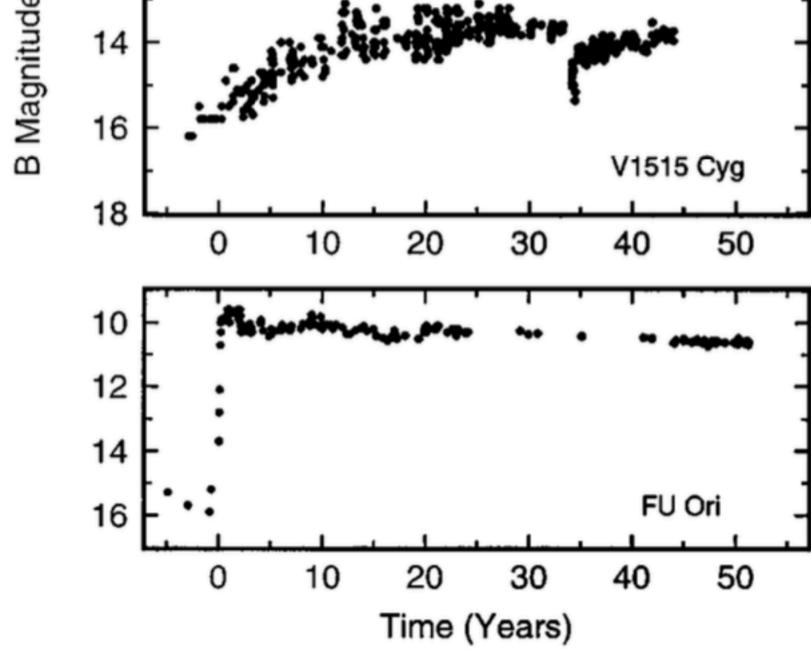




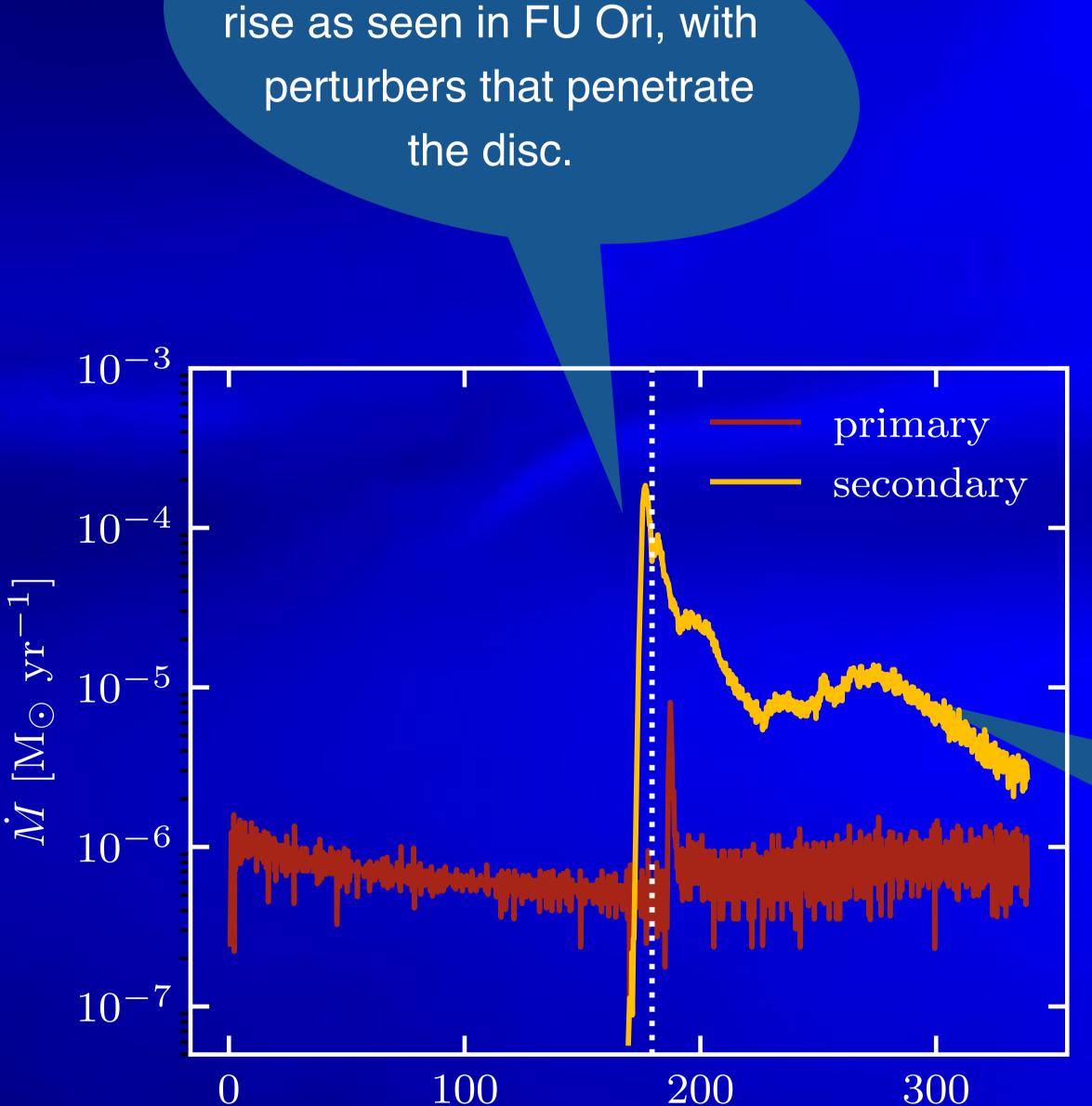








In 1936, a previously unremarkable 16th magnitude star in Orion brightened by 6 magnitudes and remained bright ever since. [1] This star is called FU Ori (bottom panel)



FU Ori N FU Ori S

FU Ori is a binary system where the low mass star (top source) was discovered to be in outburst. [2]

> Mass accretion rate of the secondary continues at a higher level to what it had been pre-flyby for the primary





On the rise times in FU Orionis events

Elisabeth M. A. BORCHERT<sup>1</sup>, Daniel J. PRICE<sup>1</sup>, Christophe PINTE<sup>1,2</sup>, Nicolás CUELLO<sup>2</sup>



[1] Hartmann L., Kenyon S. J., 1996, ARA&A , 34, 207

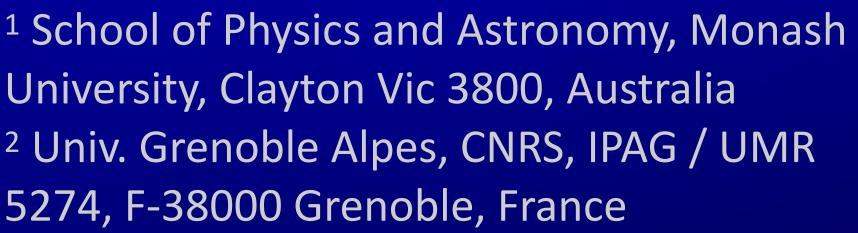
[2] Cuello N., Ménard F., Price D. J., 2023, European Physical Journal Plus



**3D SPH simulations using PHANTOM with live** 

radiative transfer through MCFOST of an unbound stellar flyby on a parabolic orbit.





**Australian Government**